# 1 Wobblers and Rayleigh-Taylor Instability Mitigation in HIF Target Implosion 2 S. Kawata<sup>a</sup>, T. Kurosaki<sup>a</sup>, K. Noguchi<sup>a</sup>, T. Suzuki<sup>a</sup>, S. Koseki<sup>a</sup>, D. Barada<sup>a</sup>, Y. Y. Ma<sup>a</sup>, A. I. 3 Ogoyski<sup>b</sup>, J. J. Barnard<sup>c</sup> and B. G. Logan<sup>c</sup> 4 5 <sup>a</sup>Utsunomiya University, 7-1-2 Yohtoh, Utsunomiya 321-8585, Japan 6 <sup>b</sup>Varna Technical University, Varna 9010, Bulgaria 7 8 <sup>c</sup>Lawrence Berkeley National Lab. and Virtual National Lab. for Heavy Ion Fusion, Berkeley, CA 9 94720, USA 10 11 **ABSTRACT** 12 A few % wobbling-beam illumination nonuniformity is realized in heavy ion inertial 13 confinement fusion (HIF) by a spiraling beam axis motion in the paper. So far the wobbling heavy ion beam (HIB) illumination was proposed to realize a uniform implosion in HIF. However, the 14 15 initial imprint of the wobbling HIBs was a serious problem and introduces a large unacceptable energy deposition nonuniformity. In the wobbling HIBs illumination, the illumination 16 17 nonuniformity oscillates in time and space. The oscillating-HIB energy deposition may contribute 18 to the reduction of the HIBs' illumination nonuniformity and also the mitigation of the Rayleigh-19 Taylor instability. The wobbling HIBs can be generated in HIB accelerators and the oscillating frequency may be several 100MHz~ 1GHz. Three-dimensional HIBs illumination computations 20 21 presented here show that the few % wobbling HIBs illumination nonuniformity oscillates 22 successfully with the same wobbling HIBs frequency. 23 24 Keywords: Heavy Ion Fusion 25 Beam Illumination Uniformity 26 27 Rayleigh-Taylor Instability Mitigation Wobbling Heavy Ion Beam 28

29

#### 1. Introduction

Heavy ion beams (HIB) have preferable features in inertial confinement fusion (ICF) [1], high energy density physics [2-4], also ion cancer therapy, etc: a HIB pulse shape is controlled precisely to fit various requirements, a HIB axis is also precisely controllable, a HIB generation energy efficiency is 30~40%, a HIB particle energy deposition is almost classical, and the deposition profile is well defined. Especially HIBs deposit their main energy in a deep area of the target material.

One of important issues in ICF is the fuel target implosion uniformity; a sufficiently uniform driver energy deposition is required [1,5,6]. Therefore, so far a dynamic stabilization method for the Rayleigh-Taylor (R-T) instability in the target implosion has been proposed and studied [7-9]. On the other hand, the HIB axis controllability provides a unique tool to smooth the HIB energy deposition nonuniformity, and can introduce wobbling or axis-oscillating HIBs [10,11]. The wobbling HIBs may contribute to reduce the HIBs' illumination nonuniformity. In addition, our previous work showed that the R-T instability growth can be reduced by the sinusoidally oscillating acceleration in time and space [12,13], if the wobbling HIBs provides the oscillating acceleration required and at the same time the wobblers' illumination is sufficiently uniform. In this sense, the wobbling HIBs may have a potential to realize a uniform energy deposition. So far the wobbling HIB has been studied to smooth the HIB illumination for a cylindrical target [3,14].

Detailed studies of three-dimensional HIBs illumination show that HIBs illumination uniformity depends strongly on the HIBs illumination scheme [15,16]. For the wobbling HIBs illumination (see Figs. 1 and 2), Ref. 16 showed a sufficiently small nonuniformity for circularly wobbling HIBs in a steady state, and the nonuniformity was evaluated after the wobbling HIBs energy deposition becomes steady [16]. However, the wobbling HIBs illumination is time-dependent. We found that in particular the initial HIBs illumination nonuniformity (the initial imprint) becomes large and is not acceptable in ICF. In this paper we present another HIBs illumination scheme for a direct drive

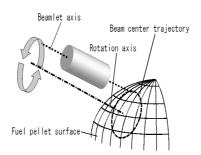


Fig. 1 The schematic diagram of wobbling HIB illumination.

spherical target. In the HIBs illumination scheme proposed here, each HIB axis has a spiraling

trajectory (see Fig. 3), so that the time-dependent HIBs illumination realizes a sufficiently low

nonuniformity (<3.57%) from the initial time to the HIBs pulse termination.

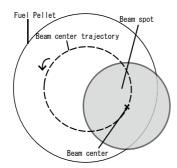


Fig. 2 A wobbling beam illumination on a direct-driven spherical target.

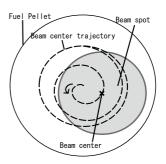


Fig. 3 Spiraling HIB. The Spiraling beam axis motion provides a sufficiently low HIBs illumination nonuniformity during the HIBs pulse duration.

# 2. Uniformity evaluation method

In our studies we employ Pb<sup>+</sup> ion HIBs with the mean particle energy of 8 GeV. The beam radius at the entrance of a reactor chamber wall  $R_{en}$  is 35 mm (see Fig. 4), the reactor chamber radius  $R_{ch}$  is 3 m. The beam particle density distribution is the Gaussian one. The longitudinal temperature of HIB ions is 100 MeV with the Maxwell distribution. The beam transverse emittance is 3.2mm mrad, from which the focal spot radius  $R_f$  and f (see Fig. 4) are obtained. The target temperature increases linearly during the time of a HIB pulse deposition from 0.025 eV to 300 eV in our study. We employ an Al monolayer pellet target structure with a 4.0 mm external radius. In our study of the HIBs illumination uniformity, we use the OK code [17], in which the detailed ion energy stopping power is computed including the beam temperature. The 32-HIBs rotation axes positions are given as presented in Ref. 18. The HIBs illumination nonuniformity is evaluated by the global rms, including also the Bragg peak effect in the energy deposition profile in the target radial direction [15]. The mode analyses are also performed to find the dominant mode of the illumination nonuniformity [15]. The stopping power of a target is the sum of the energy deposited in target nuclei, target bound and free electrons, and target ions [19-21].

In HIB ICF the Bragg peak area of the HIB energy deposition is most important for target implosion. We employ the relative *rms* nonuniformity including the Bragg peak effect, that is, the energy deposition profile:  $\sigma_i^{rms} = \frac{1}{\langle E \rangle_i} \sqrt{\sum_j^{n_\theta} \sum_k^{n_\phi} \left( \langle E \rangle_i - E_{i,j,k} \right)^2 / n_\theta n_\phi}$ ,  $\langle \sigma_{rms} \rangle = \sum_i^{n_r} w_i \sigma_i^{rms}$ ,

 $w_i = E_i/E$ . In this paper each HIB is divided into many beamlets and the spherical target Al shell is

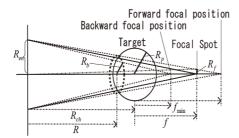


Fig. 4 A HIB illuminating on a fuel target. Pb<sup>+</sup> ion HIBs are employed with the mean particle energy of 8 GeV. The beam radius at the entrance of a reactor chamber wall  $R_{en}$  is 35 mm, and the reactor chamber radius  $R_{ch}$  is 5 m;  $R_b$  is the beam radius on a target.

- divided by fine meshes in the radial *i*-direction, the polar angle *j*-direction and the azimuthal angle
- 2 k-direction [20,22]. Here  $\langle \sigma_{rms} \rangle$  is the global rms nonuniformity,  $\sigma_i^{rms}$  is the rms nonuniformity on
- 3 the *i*-th surface of deposition,  $w_i$  is the weight function in order to include the Bragg peak effect of
- 4 the deposition profile,  $n_r$ ,  $n_\theta$  and  $n_\phi$  are the total mesh numbers in each direction of the spherical
- 5 coordinate,  $\langle E \rangle_i$  is the mean deposition energy on the *i*-th surface,  $E_i$  is the total deposition energy
- 6 in the *i*-th shell, and *E* is the total deposition energy.

We also performed mode analyses by the spherical harmonic function  $Y_n^m(\theta, \phi)$  with the amplitude of  $S_n^m$  of energy spectrum, n and m are the mode numbers, and  $\theta$  and  $\phi$  are the polar and azimuthal angles, respectively.  $E(\theta, \phi)$  is the HIB deposition energy summarized over the radial direction at each  $(\theta, \phi)$  mesh point of a target. The summation of the energy spectrum amplitude is normalized to be 1.0 in our study.

## 3. Spiraling heavy ion beam illumination uniformity

So far the wobbling HIBs illumination on a spherical fuel target has been considered for the time averaged illumination [16]. However, it was found in this paper that the initial imprint of the rotating HIBs is serious and introduces a large illumination nonuniformity. In order to reduce the HIBs illumination nonuniformity from the initial time to the beam pulse end, in this paper we propose that each HIB axis has a spiral trajectory as shown in Fig. 3 especially in the initial few rotations. After the initial few spiral rotations, the HIB axes trajectories approach to circles.

In this present work the HIBs axes alignment in Ref. 18 is employed; each HIB rotation center is defined by Ref. 18. In addition to the axes alignment, the rotation starting point of each HIB must be specified. The 32-HIBs illumination scheme consists of 6 circles of latitude and the two HIBs illuminating from the top and bottom. Each circle of latitude has 5 HIBs. HIBs located at the even

number of the circle of latitude start from the top position of the rotation circle and the HIBs in the odd number of the rings start from the bottom of the circle.

When we do not employ the spiraling trajectory as shown in Fig. 2, the illumination nonuniformity history shows an unacceptable large nonuniformity during the first few rotations as presented in Fig. 5 (the solid line). In this case, the direct drive target radius is 4mm, the energy deposition layer consists of Al, and we employ  $Pb^+$  ion HIBs with the mean energy 8GeV, as shown above. The HIB axis rotation radius is 2mm and the HIB radius is 3mm. When the spiraling HIBs are employed for the first two rotations, the HIBs illumination uniformity is drastically improved as shown in Fig. 5 (the dotted curve). In Fig. 5 the time is normalized by the wobbling beam axis rotation time  $\tau$ .

Figure 6 shows the history of the spiraling HIBs illumination nonuniformity as well as the HIB particle illumination loss history. Here, we define the loss as the percentage of particles that do not hit the target. By the spiraling trajectory of the HIB axis, a sufficiently small nonuniformity is successfully realized, and a small part of the HIBs particles do not hit the target in order to obtain

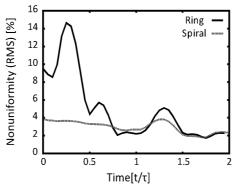


Fig. 5 Histories of energy deposition nonuniformities for the circulary moving axies (solid line) and for the spiraling HIBs. The spiraling HIB-axis motion realizes a rather low HIBs illumination nonuniformity.

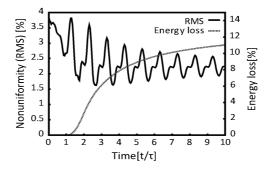


Fig. 6 Histories of the spiraling HIBs illumination nonuniformity (the solid line) and the illumination loss (the dotted line). The spiraling HIBs provide a sufficiently uniform deposition uniformity in HIF.

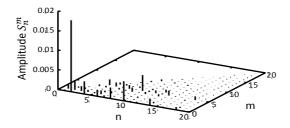


Fig. 7 The energy spectrum at  $t = 1.3\tau$  of the spiraling HIBs. The mode (n, m)=(2, 0) is dominant throughout the spiral HIBs illumination.

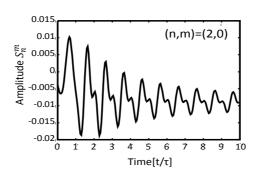


Fig. 8 The mode (2, 0) amplitude of HIBs energy deposition nonuniformity versus time.

the illumination uniformity (see Fig. 6 (the dotted line)). Figure 7 shows an energy spectrum at  $t = 1.3\tau$ , at which time the HIBs illumination nonuniformity has a local-peak value as shown in Fig. 5 (the black line). In Fig. 7, (n, m) are the polar and azimuthal mode numbers, and  $S_n^m$  is the amplitude of the spectrum, respectively. If the deposition energy distributed is prefectly spherically symmetric, the amplitude of the spectrum is 1.0 in the mode (n, m) = (0, 0) in our study. For this reason, the amplitude of the mode (n, m) = (0, 0) becomes large, nearly 1.0. In this paper the amplitude of the spectrum the mode (n, m) = (0, 0) is not displayed. As a result, the amplitude of spectrum mode (n, m) = (2, 0) is the largest mode in Fig. 7, and the mode (n, m) = (2, 0) is dominant throughout the HIBs illumination. Figure 8 shows the amplitude of the mode (2, 0) versus time, and

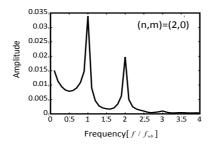


Fig. 9 Spectrum of the mode (2, 0) in its frequency space. The frequency  $f_{\rm wb}$  is the wobbling HIB frequency. The small nonuniformity of the HIBs energy deposition has the same frequency  $f_{\rm wb}$  of the wobblers and also

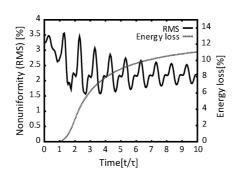


Fig. 10 Histories of the spiraling HIBs illumination nonuniformity (the solid line) and the illumination loss (the dotted line). The spiraling HIB radius changes at t=1.3t from the initial beam radius of 3.1mm to

Fig. 9 presents the spectrum of the mode (2, 0) amplitude in its frequency space. In Fig. 9  $f_{wb}$  shows the wobbling HIBs rotation frequency. When each HIB deposits its energy on a fuel target with the function of  $e_i(\vec{r},t)(1+\delta_i(\vec{r})e^{i\Omega t})$ , the total HIBs energy deposition  $E(\vec{r},t)$  can be expressed as  $E(\vec{r},t) = \sum_{i=1}^{N_b} e_i(\vec{r},t)(1+\delta_i(\vec{r})e^{i\Omega t}) = \sum_{i=1}^{N_b} e_i(\vec{r},t) + \{\sum_{i=1}^{N_b} e_i(\vec{r},t)\delta_i(\vec{r})\}e^{i\Omega t}$ . Here  $\delta_i(\vec{r})e^{i\Omega t}$  is the oscillating part of the deposited energy,  $\Omega = 2\pi f_{wb}$ ,  $\delta_i \ll 1$ ,  $N_b$  is the total HIBs number, and in our case  $N_b$ =32. Therefore, the wobbling beams induce naturally the deposition energy oscillation with the wobbling oscillation frequency of  $\Omega$ .

The result in Fig. 9 demonstrates that the small nonuniformity of the HIBs energy deposition has the oscillation with the same frequency and the double frequency with the wobbling HIBs oscillation frequency of  $f_{wb}$ . In addition, the total nonuniformity magnitude is suppressed less than 3.87%. The results in this paper show a possibility of a rather uniform implosion in heavy ion fusion based on the wobbling HIBs. When a perturbation inducing the implosion nonuniformity emerges during the fuel target implosion based on the origin of the deposition energy nonuniformity, the perturbation may grow from the energy deposition nonuniformity with a certain phase. In the wobbling HIBs illumination, the nonuniformity phase is defined by the wobbling beams' motion. The overall perturbation growth is the superposition of all the perturbations with the different initial phases but with the same wavelength. Based on these considerations, it would be pointed out that the wobbling HIBs may contribute to a fuel target uniform implosion in heavy ion inertial fusion [12,13,22].

In the analyses presented above, the HIB radius was fixed to be 3mm. In the wobbling HIBs illumination, the peak of the HIB energy deposition nonuniformity appears in the first few rotations. When the beam radius changes from 3.1mm to 3mm at  $t = 1.3\tau$  during the spiral motion in the second rotations, we have an additional improvement for the HIBs illumination nonuniformity. The peak value of the nonuniformity becomes less than 3.57% (see Fig. 10), although the peak value of the nonuniformity in Fig. 6 was about 3.87%.

### 4. Robustness of HIB illumination scheme to target alignment

In this section a robustness of the spiraling HIBs' illumination scheme is examined against the target alignment error in a fusion reactor. We simulate the effect of a little displacement dz as well as dx and dy on the HIB illumination nonuniformity as shown in Fig. 11. Figure 12 shows the relation between the pellet displacement and the maximal rms nonuniformity in the best case, in which each HIB radius changes from 3.1 mm to 3.0 mm at  $t=1.3\tau$ . The maximal rms nonuniformity is less than 4.5 % for the pellet displacement less than about 80  $\mu$ m. The HIB particle illumination loss (mishit) is about 11%, and it is not so serious up to 90 $\mu$ m. Figure 13 presents the spectrum of the mode (2, 0) amplitude in its frequency space. The result shown in Fig. 13 shows that the target alignment error in a reactor does not influence the oscillation frequency of the small HIBs illumination nonuniformity. The nonuniformity oscillates again with the wobbling frequency.

In the conventional beam illumination scheme, a fuel target displacement of  $50-100 \mu m$  is tolerable [15]. The result shown in Fig. 12 means that the spiraling HIBs' illumination scheme

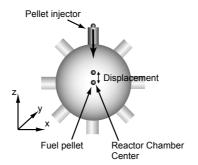


Fig. 11 Traget alignment error in a fusion reactor.

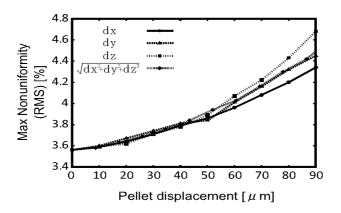


Fig. 12 The relation between the pellet displacement and the maximal rms nonuniformity in the best case, in which each HIB radius changes from 3.1 mm to 3.0 mm at  $t=1.3\tau$ . The maximal rms nonuniformity is less than 4.5% for the pellet displacement less than about 80  $\mu$ m.

- 1 provides the same order of the tolerable displacement with the conventional one. An enhancement
- 2 of the allowable range should be studied further in the future, if required.

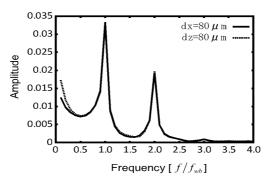


Fig. 13 The spectrum of the mode (2, 0) amplitude in its frequency space for dx=80mm and for dz=80mm.

# 5. Dynamic mitigation of Rayleigh-Taylor instability

In general a perturbation of physical quantity would be an origin of instability. Normally the perturbation phase is unknown so that the instability growth is discussed with the growth rate. However, if the perturbation phase is known, the instability growth can be controlled by a superposition of perturbations (see Fig. 14); the most well-know mechanism is a feedback control to compensate the displacement or the distortion of physical quantity. If the perturbation is induced by, for example, a particle beam axis oscillation or wobbling, the perturbation phase could be controlled and the instability growth is mitigated by the superposition of the growing perturbations. In plasmas it is difficult to detect the perturbation phase and amplitude. However, in plasmas, if we can actively impose the perturbation phase by the driving energy source wobbling or so, and therefore, if we know the phase of the perturbations, the perturbation growth can be controlled in a similar way [12, 13, 22]. This control mechanism is apparently different from the dynamic stabilization shown in the previous works [9]. The wobbling heavy ion beams define the perturbation phase. This means that the perturbation phase is known, and so successively imposed perturbations are superposed on plasma. We can use the capability to reduce the instability growth by the phase-controlled superposition of perturbations.

If the HIB wobbles uniformly in time, the imposed perturbation for a physical quantity of F at  $t=\tau$  may be written as  $F=\delta Fe^{i\Omega\tau}e^{\gamma(t-\tau)+i\vec{k}\cdot\vec{x}}$ . Here  $\delta F$  is the amplitude,  $\Omega$  the wobbling or oscillation frequency, and  $\Omega\tau$  the phase shift of superposed perturbations. At each time  $t=\tau$ , the wobbler provides a new perturbation with the controlled phase shifted and amplitude defined by the driving wobbler itself. After the superposition of the perturbations, the overall perturbation is described as

1 
$$\int_0^t d\tau \ \delta F e^{i\Omega\tau} e^{\gamma(t-\tau) + i\vec{k}\cdot\vec{x}} \propto \frac{\gamma + i\Omega}{\gamma^2 + \Omega^2} \delta F e^{\gamma t} e^{i\vec{k}\cdot\vec{x}}. \tag{1}$$

At each time of  $t = \tau$  the driving wobbler provides a new perturbation with the shifted phase. Then each perturbation grows by the factor of  $e^{\gamma t}$ . At  $t > \tau$  the superposed overall perturbation growth is modified as shown above. When  $\Omega \gg \gamma$ , the perturbation amplitude is reduced by the factor of  $\frac{\gamma}{\Omega}$ , compared with the pure instability growth ( $\Omega = 0$ ) based on the energy deposition nonuniformity. The example simulation results also support the effect of the dynamic mitigation mechanism well [13].

From the analytical expression for the physical quantity F in Eq. (1), the mechanism proposed in this paper does not work, when  $\Omega \ll \gamma$ . Only for the modes, which satisfy the condition of  $\Omega \geq \gamma$ , the mechanism of the instability mitigation by the wobbler can be applied for its growth mitigation. For R-T instability, the growth rate  $\gamma$  tends to become larger for a short wavelength. If  $\Omega \ll \gamma$ , the modes cannot be mitigated. In addition, if there are other sources of perturbations in the physical system and if the perturbation phase and amplitude are not controlled, this dynamic mitigation mechanism also does not work. For example, if the sphericity of an inertial fusion fuel target is degraded, the dynamic mitigation mechanism does not work. In this sense the dynamic mitigation mechanism is not almighty. Especially for a uniform compression of an inertial fusion fuel all the instability stabilization and mitigation mechanisms would contribute to release the fusion energy in stable.

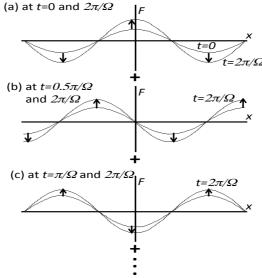


Fig. 14 Superposition of perturbations defined by the wobbling HIBs. At each time the wobbler provides a perturbation, whose amplitude and phase are defined by the wobbler itself. If the system is unstable, each perturbation is a source of instability. At a certain time the overall perturbation is the superposition of the growing perturbations. The superposed perturbation growth is mitigated by the beam wobbling motion.

#### 6. Conclusions

In this paper, we presented the HIBs illumination scheme of the spiraling HIBs and that by the spiraling HIBs the HIBs illumination uniformity is drastically improved. Throughout the HIB input pulse, the beam illumination nonuniformity is kept low, that is, less than about 3.57%. The sufficiently small illumination nonuniformity is successfully realized by the wobbling HIBs onto a spherical target. The small-amplitude nonuniformity oscillation frequency is mostly the wobbling oscillation frequency and also the double of wobbling frequency. The wobbling HIBs may supply a viable uniform implosion mode in heavy ion fusion.

The oscillating small nonuniformity of the HIBs deposited energy may induce an oscillating acceleration, which reduces significantly the R-T instability growth [12, 13, 22]. Therefore, the spiraling HIBs may provide a stable implosion of a spherical inertial fuel target in the direct-driven scheme.

When the HIB input pulse is long so that HIBs radius zooming is needed [23], we may also employ the inversely-spiraling illumination scheme to realize the zooming HIBs illumination to implode a converging compressed fuel with a low illumination nonuniformity. In addition, recently A. Friedman has proposed an interesting idea of the arc-based wobbling HIB smoothing [24]. In contrast to the full-rotation approach presented in this paper, the arc-based wobbling HIBs provide another promising approach to smooth the HIBs illumination nonuniformity in high-energy density physics and in heavy ion fusion target implosion.

### Acknowledgments

This work is partly supported by MEXT, JSPS, ILE/Osaka Univ., KEK and CORE (Center for Optical Research and Education, Utsunomiya university, Japan). The works is also partly supported by the Japan/U.S. Fusion Research Program.

#### References

- [1] S. Atzeni, J. Meyer-Ter-Vehn, "The Physics of Inertial Fusion", Oxford Science Pub., 2004.
- 29 [2] National Res. Council of the National Academies, "Frontiers in High Energy Density Physics", the National Academies Press, Washington, D. C., U.S.A. 2003.
- 31 [3] N. A. Tahir, et al., Th. Plasma Phys. Control. Fusion 53 (2011) 124004.

22

- 1 [4] N. A. Tahir, et al., Phys. Rev. Lett. 95 (2005) 035001.
- 2 [5] M. H. Emery, et al., J. P. Phys. Rev. Lett. 48 (1982) 253.
- 3 [6] S. Kawata, et al., J. Phy. Soc. Jpn. 53 (1984) 3416.
- 4 [7] F. Troyon, Phys. Fluids 14 (1971) 2069.
- 5 [8] J. P. Boris, Comments Plasma Phys. Cont. Fusion **3**, 1 (1977).
- 6 [9] R. Betti, et al., Phys. Rev. Lett. 71 (1993) 19.
- 7 [10] R. C. Arnold, et al., Nucl. Inst. Meth. Phys. 199 (1982) 557.
- 8 [11] H. Qin, et al., Phys. Rev. Lett. 104 (2010) 254801.
- 9 [12] S. Kawata, et al., Laser Part. Beams 11 (1993) 757.
- 10 [13] S. Kawata, et al., Nucl. Instr. Meth. Phys. Res., A606 (2009) 152.
- 11 [14] M. M. Basko, et al., J. Phys. Plasmas 11 (2004) 1577.
- 12 [15] S. Miyazawa, et al., Phys. Plasmas 12 (2005) 122702.
- 13 [16] J. Runge, et al., Phys. Plasmas 16 (2009) 133109.
- 14 [17] A. I. Ogoyski, et al., Comput. Phys. Comm. 181 (2010) 1332 and references therein.
- 15 [18] S. Skupsky, et al., J. Appl. Phys. 54 (1983) 3662.
- 16 [19] T. A. Mehlhorn, J. Appl. Phys. 52 (1981) 6522.
- 17 [20] T. Peter, T., J. Meyer-ter-Vehn, Phys. Rev. A 43 (1991) 1998.
- 18 [21] T. Peter, J. Meyer-ter-Vehn, Phys. Rev. A 43 (1991) 2015.
- 19 [22] S. Kawata, Phys. Plasmas 19 (2012) 024503.
- 20 [23] B. G. Logan, et al., J. J. Phys. Plasmas 15 (2008) 072701.
- 21 [24] A. Friedman, Phys. Plasmas 19 (2012)063111.